# Machine Learning Applications in Dark Matter Search



### **Previous work in High Energy Physics**

Improving the significance of resonance signal (K\*0) using Machine Learning Techniques in high energy heavy-ion collisions. Invariant mass distribution: ALICE data



Ref. chap 3 and Appendix of thesis

#### **Introduction to Dark Matter**

- Compelling evidence from a number of astrophysical observations.
- Hypothetical form of matter thought to account for ~27% of the energy density of the universe.
- Does not interact with electromagnetic radiation, hence called "dark".
- Non-baryonic in nature and possibly composed of yet undiscovered particles e.g Weakly interacting massive particles (WIMPs).
- They interact with gravity and other weak-like forces which are not a part of the Standard Model.
- We will be studying the dark matter detection experiment of SuperCDMS.

#### **Dark Matter Search by SuperCDMS**

- **Experiment**: Super Cryogenic Dark Matter Search (**SuperCDMS**).
- **Search**: WIMPs as a dark matter candidate.
- **Measures**: Phonon and charge signals.
- Signal: Nuclear Recoils (NR).
- **Background**: Electron Recoils (ER).
- My work: Separate ER and NR using Machine Learning Techniques.



#### **Flow Chart of Analysis**



#### Data Set: Model

- Events were taken from <sup>252</sup>Cf<sub>98</sub> and <sup>133</sup>Ba<sub>56</sub> data.
- As dark matter particles have not been discovered yet, we need a system to mimic the dark matter signal, and a system to mimic the background events.
- Dark matter signal: Neutrons from Cf source (Nuclear recoils)
- Background: Gamma from Ba source. (Electron recoils)

#### **Data and Preselection cuts**

# Data was collected during the period Aug 2013-Aug 2014.

**Not Empty** Removes empty events in the detector with no trigger.

- **Good Flashtime** Removes events when the detector was not flashed for period greater than 3300 seconds.
- **Base Temperature** Removes events when the base temperature was not in the good range.
- **Voltage Bias** Ensures voltage was maintained within 4V range.
- **Good Start time** Removes events with Non Stationary Optimal Filter delay.

Analysis Phonon recoil energy greater than 10 keV.

threshold

### WIMP Signal Model



- Blue is WIMP spectra from theory.
- Green is Cf data distribution.
- Red is detector efficiency.
- Black is the corrected spectrum.



- Weight vector is calculated by taking the ratio of corrected sprectrum (black) and Cf data (green)
- Weigh each event of the Cf data by the corresponding weight in the weight vector.

#### **FV: Phonon Radial Cut**

Z1: Barium-ba, prpart vs ptNF T2Z1: Barium, Weeks1-6, prpartOF, Bin No. 10, 1.00 <ptNF < 1.10  $\gamma^2$  / ndf = 131.6 / 36 **Projection along Y for** 0.5 stil 000  $934.9 \pm 13.6$ Constant Data Gaussian Fit a ptNF bin, fitted with  $\frac{1}{2}$   $\mu$  + 2 $\sigma$  fit of Projection-Y 0 2495 ± 0 0004 -Phonon Radial Cut: Exponential Fit 0.45 a gaussian function Sigma 0.03433 ± 0.00032 800 0.4 0.35 600 prpartOF 6.0 400 0.25 0.2 200 0.15 81 0.3 0.15 0.2 0.25 0.35 0. 0.4 0.45 0.5 12 10 14 prpartOF ptNF [keV]

- prpartOF is the ratio of outer phonon sum and total phonon energy. ptNF is total phonon energy
- Phonon radial cut removes high-radius phonon events i.e events above the Exponential fit.

#### **FV: Charge Symmetric Cut**

 Inner charges collected from side 1 and side 2.

 Geometry of the detector→
Asymmetric charge collection near surface, symmetric charge collection in the bulk region. Inner Charge (S2): qi2OF 9 & 0 Charge Symmetric cut -2 2 10 Inner Charge (S1): qi1OF

Z1: Barium-ba, Weeks 1-6: Inner Charge(S2) vs Inner Charge(S1)

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#### **FV: Charge Radial Cut**

Z1: Barium-ba, Side 1: Outer Charge vs Inner Charge

Z1: Barium-ba, Side 2: Outer Charge vs Inner Charge

5 Outer Charge **Duter Charge** Charge Radial Cut Charge radial cut З 2 C -2Ľ -2 10 2 6 8 0 2 8 10 0 6 Inner Charge Inner Charge

Removes events closer to the cylindrical sidewall of the detector i.e larger value of outer charges.

#### **ML classification of NR and ER events**

• We have used Cf data as a source of NR events and Ba data as a source of ER events.

#### **Features used for Training our BDT**

precoiltNF:	the non-luke phonon energy
qsummaxOF:	sum for the side with the maximum qsum (inner+outer charge)
ytNF:	Yield i.e qsummaxOF/precoiltNF
qzpartOF:	Z-direction fiducializing parameter for charge
qrpartOF:	Radial fiducializing parameter for charge
pzpartOF:	Z-direction fiducializing parameter for phonon
prpartOF:	Radial fiducializing parameter for phonon

#### Input features used for training of BDT



### **BDT distributions**



- We take the BDT cut value to be -0.0528
- Events with BDT response> BDT cut value are classified as signal (NR).

#### **ML Application Phase results: ER NR bands**



## **ML Application Phase results: Yield bands**





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#### FV cuts and BDT cut



#### FV cuts and BDT cut: yield bands



# Conclusion

- BDT cut provides an event by event separation of Electron Recoils and Nuclear Recoils.
- We achieve an enhanced NR band after FV cuts and BDT classification.
- We have demonstrated a way to get a pure NR band (and remove ER background contributions) in the detector using ML techniques, this can be adapted for DM search.

# Outlook

 Further classification of the NR events which may have possibly come from WIMPs and neutrons.

# Backup

#### **SuperCDMS experiment detector**

- The Super Cryogenic Dark Matter Search (**SuperCDMS**) experiment is designed to directly detect particle dark matter in the form of WIMPs.
- The Ge-made interleaved Z-sensitive lonization and Phonon (iZIP) detectors aims at measuring and distinguishing ionization and phonon signals for nuclear recoils(NR) produced by WIMPs and electron recoils(ER) produced by background sources.



Figure reference: *Applied Physics Letters*, vol. 103, no. 16, p. 164105, 2013.



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Figure reference: https://thesis.library.caltech.edu/11056/

#### **Fiducial Volume Optimization**

- Fiducial Volume Optimization→ removes unwanted surface events and outer edge events. ("Fiducialization")
- These unwanted events lead to misidentification of ERs with suppressed ionization collection as NRs.
- Fiducial Volume (FV) of the detector → Hypervolume which is set by optimizing cuts in various parameter space.
- Events inside the FV are called "bulk" events and events outside the FV are called "surface" events.

#### Features chosen (Reasons)

precoiltNF:	Difference in distribution at low phonon energy
qsummaxOF:	sum for the side with the maximum qsum (inner+outer charge)
ytNF:	Yield i.e qsummaxOF/precoiltNF, a classic CDMS discriminator.
qzpartOF:	Z-direction fiducializing parameter for charge, removes near surface events
qrpartOF:	Radial fiducializing parameter for charge
pzpartOF:	Z-direction fiducializing parameter for phonon
prpartOF:	Radial fiducializing parameter for phonon

#### **Theory spectra from Lewin and Smith**

$$\frac{\mathbf{d}\mathbf{R}}{\mathbf{d}\mathbf{E}_{\mathbf{R}}} = \frac{\mathbf{R}_{\mathbf{0}}}{\mathbf{E}_{\mathbf{0}}\mathbf{r}}\mathbf{e}^{-\mathbf{E}_{\mathbf{R}}/\mathbf{E}_{\mathbf{0}}\mathbf{r}}$$

- $\rightarrow~{\bf E_R}$  is the Recoil Energy
- $\rightarrow~E_0$  is the most probably KE of dark matter of mass  $\mathbf{M_D}$
- $\rightarrow~\mathbf{R}$  is the event rate per unit mass
- $\rightarrow~\mathbf{R_0}$  is the total event rate

 $\rightarrow \mathbf{r} = 4\mathbf{M_D}\mathbf{M_T}/(\mathbf{M_D} + \mathbf{M_T})^2$ 

- Event rate is expressed in events/keV\*kg\*day or "dru".
- Reference: J.D. L.ewin, PE Smith/Astroparticle Physics 6 (1996)

#### **Final Numbers after FV and BDT cut**

- Passage fraction (ratio of number of events passing the FV cut and total number of events)=67.7 %
- The fraction of events classified as NR by ML improves from 54.49 % to 56.28 % after the application of FV and BDT cut.

#### **BDT algorithm flow chart**



Based on the output of all trees in the forest, each event is assigned a BDT response valu ranging from -1 to +1. Background-like events will have a BDT response value shifted towards -1, and signal-like events will have a response value shifted towards +1

#### **Adaptive Boosted Decision Trees**

## AdaBoost: learning ensemble

 $\hat{\mathbf{w}}_{t} = \frac{1}{2} \ln \left( \frac{1 - weighted\_error(f_{t})}{weighted\_error(f_{t})} \right)$ • Start same weight for all points:  $\alpha_i = 1/N$ • For t = 1,...,T  $\boldsymbol{\alpha}_{i} \leftarrow \begin{bmatrix} \boldsymbol{\alpha}_{i} e^{-W_{t}}, & \text{if } f_{t}(\mathbf{x}_{i}) = \mathbf{y}_{i} \\ \boldsymbol{\alpha}_{i} e^{\hat{W}_{t}}, & \text{if } f_{t}(\mathbf{x}_{i}) \neq \mathbf{y}_{i} \end{bmatrix}$ - Learn  $f_t(\mathbf{x})$  with data weights  $\alpha_i$ Compute coefficient  $\hat{w}_{t}$ - Recompute weights  $\alpha_i$ - Normalize weights  $\alpha_i$  Final model predicts by:  $\alpha_i \leftarrow \frac{\alpha_i}{\sum_{j=1}^N \alpha_j}$  $\hat{y} = sign\left(\sum_{t=1}^{T} \hat{\mathbf{w}}_t f_t(\mathbf{x})\right)$ 

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#### **BDT** parameters

- Number of Trees = 850
- Min node size i.e Minimum percentage of training events required in a leaf node= 2.5%
- Max depth i.e max depth of DT = 3
- Learning rate for AdaBoost algorithm adaboostbeta = 0.5
- Separation Type= Gini Index
- Ncuts Number of grid points in variable range used in finding optimal cut in node splitting=20

#### **Gini Index**

#### How do I select which feature I take first?

#### **Gini Index :** Define as $p^2 + q^2$ (at a given node)

- p: fraction of positive (signal) events
- q=1-p : fraction of negative (background) events

#### Weighted Gini Split:

$$\mathbf{I_G} = \mathbf{f_{left}} * \mathbf{G_{left}} + \mathbf{f_{right}} * \mathbf{G_{right}}$$

where,

 $f_{left} =$  fraction of events which go in the left split  $f_{right} =$  fraction of events which go in the right split  $G_{left} =$  Gini index of left node  $G_{right} =$  Gini index of right node



Plot of Gini index vs p (positive/signal)

#### Correlations

#### Linear correlation coefficients in % 100 summaxOF 99 6 100 -4 80 60 prpartOF 6 2 100 40 pzpartOF 100 20 qrpart1OF -20 100 5 -0 -20 qzpartOF -13 100 -40vtNF 100 -13 -23 -60 -80 precoiltNF 100 2 2 99 -100 precoiltNF PZpartOF qs<sub>ummax</sub>OF 9rpart10F prpartOF <sup>q≥part</sup>OF YtNF

#### Correlation Matrix (signal)



Correlation Matrix (background)

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#### **Additional content**

BDT

The extracted signal of  $K^{*0}$  is fitted with Breit-Wigner distribution plus a residual background function. Mass and width for  $K^{*0}$  can be obtained from fit parameters. Breit-Wigner function is defined below:

$$\frac{Y}{2\pi} \frac{\Gamma_0}{(M_{\pi K} - M_0)^2 + \frac{\Gamma_0^2}{4}} + Res.Bkg.$$
(1)

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<sup>129</sup> where  $M_0$  and  $\Gamma_0$  are the mass and width of the K<sup>\*0</sup>.  $M_{\pi K}$  is  $\pi K$  invariant mass. The parameter <sup>130</sup> Y gives the Breit-Wigner area. The last term is residual background function, which is taken as <sup>131</sup> the polynomial of second order in invariant mass ( $AM_{\pi K}^2 + BM_{\pi K} + C$ ). The parameter Y for each <sup>132</sup>  $p_T$  bin gives the raw yield counts of K<sup>\*0</sup>. Extracted signal of K<sup>\*0</sup> is fitted with Breit-Wigner <sup>133</sup> distribution and is shown in the right panel of Fig. 14 and Fig. 15 for two different cases (mixed <sup>134</sup> background and like sign event background), where blue line indicates fit for signal with resid-<sup>135</sup> ual background and red line indicates fit for only residual background.

:	Ranking Ranking	input varia result (top	ab P	les (method specific) variable is best ranked)
1			•••	
:	Rank : V	ariable	•	Variable Importance
1			•••	
:	1 : y	tNF	;	1.721e-01
:	2 : p	recoiltNF		1.578e-01
:	3:0	summaxOF	;	1.445e-01
:	4 : p	zpartOF	;	1.425e-01
:	5 : q	zpartOF	;	1.389e-01
:	6:q	rpart10F	•	1.334e-01
:	7 : p	rpartOF	•	1.108e-01



#### Boost weights vs tree



#### **Training events selection**



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